

FI 008 – Eletrodinâmica I

1º Semestre de 2021

01/07/2021

Aula 27

Aulas passadas

Potenciais de Liénard-Wiechert:

$$\Phi(\mathbf{x}, t) = \frac{e}{|\mathbf{x} - \mathbf{r}(t_r)|} \frac{1}{1 - \frac{\dot{\mathbf{r}}(t_r) \cdot [\mathbf{x} - \mathbf{r}(t_r)]}{c|\mathbf{x} - \mathbf{r}(t_r)|}} = \frac{e}{R} \frac{1}{1 - \boldsymbol{\beta} \cdot \hat{\mathbf{n}}} \Big|_{t=t_r}$$

$$\mathbf{A}(\mathbf{x}, t) = \frac{e/c}{|\mathbf{x} - \mathbf{r}(t_r)|} \frac{\dot{\mathbf{r}}(t_r)}{1 - \frac{\dot{\mathbf{r}}(t_r) \cdot [\mathbf{x} - \mathbf{r}(t_r)]}{c|\mathbf{x} - \mathbf{r}(t_r)|}} = \frac{e}{R} \frac{\boldsymbol{\beta}}{1 - \boldsymbol{\beta} \cdot \hat{\mathbf{n}}} \Big|_{t=t_r}$$

Campos de Liénard-Wiechert:

$$\mathbf{E} = \frac{e}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{n}})^3} \left\{ \frac{\hat{\mathbf{n}} - \boldsymbol{\beta}}{\gamma^2 R^2} + \frac{1}{c} \frac{\hat{\mathbf{n}} \times [(\hat{\mathbf{n}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}]}{R} \right\} \Big|_{t=t_r}$$

$$\mathbf{B} = (\hat{\mathbf{n}} \times \mathbf{E}) \Big|_{t=t_r}$$

$$\mathbf{R} = \mathbf{x} - \mathbf{r}(t_r)$$

$$\boldsymbol{\beta} = \frac{\dot{\mathbf{r}}(t_r)}{c}$$

Notação:

$$R = |\mathbf{x} - \mathbf{r}(t_r)|$$

$$\dot{\boldsymbol{\beta}} = \frac{\ddot{\mathbf{r}}(t_r)}{c}$$

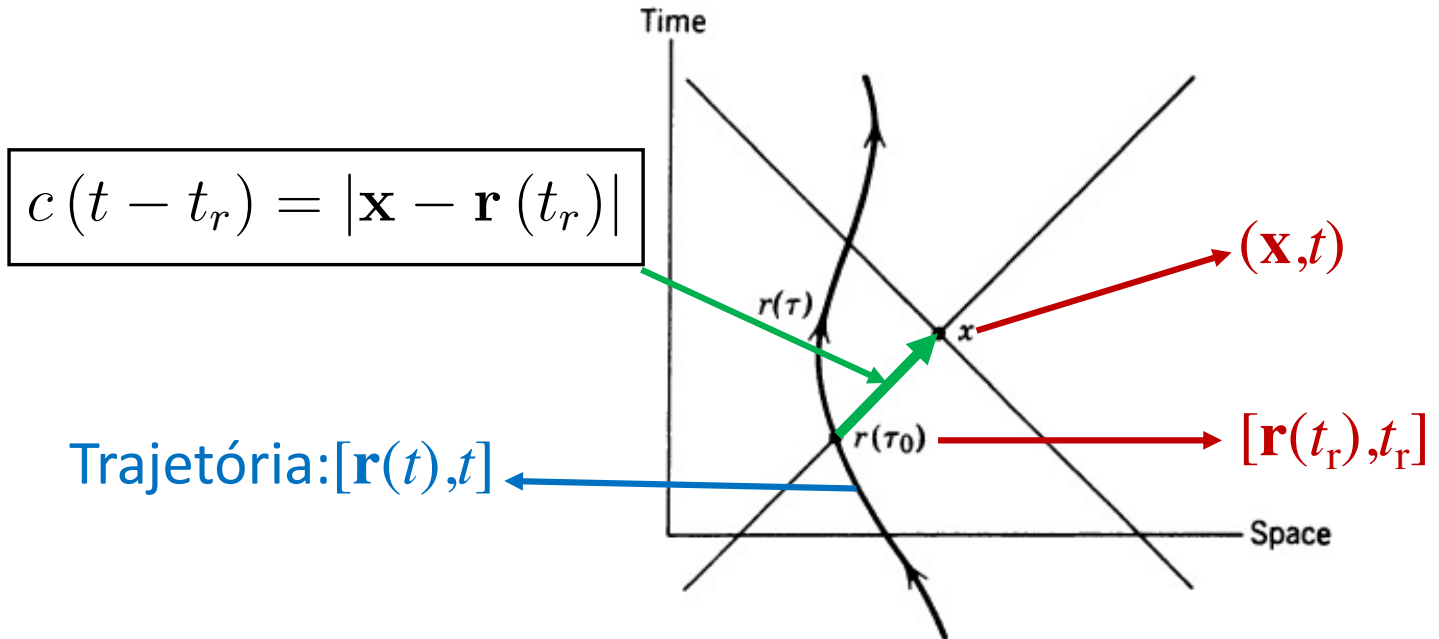
$$\hat{\mathbf{n}} = \frac{\mathbf{R}}{R} = \frac{\mathbf{x} - \mathbf{r}(t_r)}{|\mathbf{x} - \mathbf{r}(t_r)|}$$

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}}$$

Aulas passadas

O tempo retardado t_r :

(\mathbf{x}, t) está no cone de luz do futuro de $[\mathbf{r}(t_r), t_r]$



Aulas passadas

Distribuição angular da potência total irradiada:

$$\frac{dP(t_r)}{d\Omega} = \frac{e^2}{4\pi c} \frac{1}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{n}})^5} \left| \hat{\mathbf{n}} \times [(\hat{\mathbf{n}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|_{\text{ret}}^2$$
$$\frac{dP_{\text{n\~{a}o-relat.}}}{d\Omega} = \frac{e^2}{4\pi c} \left| \hat{\mathbf{n}} \times [\hat{\mathbf{n}} \times \dot{\boldsymbol{\beta}}] \right|^2 = \frac{e^2 a^2}{4\pi c^3} \sin^2 \theta$$

Potência total irradiada

$$P = \frac{2}{3} \frac{e^2}{c} \gamma^6 \left[\left| \dot{\boldsymbol{\beta}} \right|^2 - \left| \boldsymbol{\beta} \times \dot{\boldsymbol{\beta}} \right|^2 \right] \quad (\text{Fórmula de Liénard})$$
$$P_{\text{n\~{a}o-relat.}} = \frac{2}{3} \frac{e^2}{c} \left| \dot{\boldsymbol{\beta}} \right|^2 = \frac{2}{3} \frac{e^2 a^2}{c^3} \quad (\text{Fórmula de Larmor})$$

Perdas em aceleradores lineares

Da fórmula de Liénard:

$$P_{lin} = \frac{2}{3} \frac{e^2}{c} \gamma^6 \dot{\beta}^2 = \frac{2}{3} \frac{e^2}{c} \left(\frac{F}{mc} \right)^2 = \frac{2}{3} \frac{e^2}{m^2 c^3} \left(\frac{dE}{dx} \right)^2$$
$$\frac{P_{lin}}{P_{gain}} = \frac{2}{3} \frac{e^2}{m^2 c^4} \left(\frac{dE}{dx} \right) \approx \frac{dE/dx}{10^{14} \text{ MeV/m}}$$

Usando valores típicos:

$$\frac{dE}{dx} = 50 \frac{\text{MeV}}{\text{m}} \Rightarrow \frac{P_{lin}}{P_{gain}} \approx 10^{-13}$$

Perdas por radiação são desprezíveis.

Perdas em aceleradores circulares

Da fórmula de Liénard:

$$P_{circ} = \frac{2 e^2}{3 c} \gamma^4 \dot{\beta}^2 = \frac{2 e^2}{3 c} \frac{\gamma^4}{\rho^2} = \frac{2 e^4 B^2}{3 m^2 c^3} \gamma^2$$

$$\delta E_{volta} \approx \frac{4\pi e^3 B}{3 mc^2} \gamma^3$$

Usando valores típicos:

$$\left. \begin{array}{l} 10 \text{ GeV} \\ 100 \text{ m} \\ 3.3 \text{ kG} \end{array} \right\} \Rightarrow \delta E_{volta} \approx 10 \text{ MeV}$$

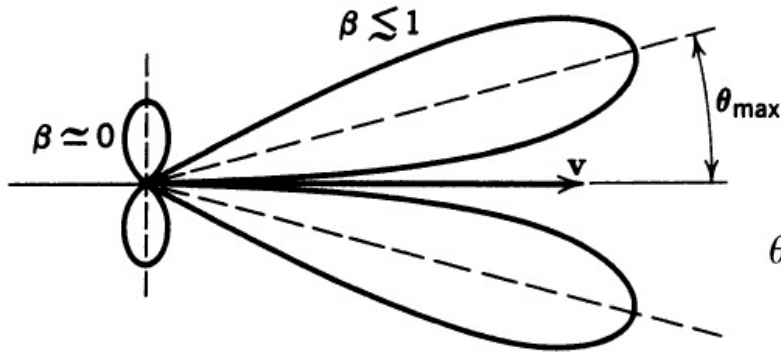
Perdas não são desprezíveis, mas da ordem do ganho de energia gerado pela radiofrequência.

Distribuição angular da radiação para velocidade e aceleração paralelas

$$\vec{\beta} \parallel \dot{\vec{\beta}}$$

$$\frac{dP(t')}{d\Omega} = \frac{e^2 \dot{v}^2}{4\pi c^3} \frac{\sin^2 \theta}{(1 - \beta \cos \theta)^5}$$

$$\gamma \gg 1 \quad \frac{dP(t')}{d\Omega} \simeq \frac{8}{\pi} \frac{e^2 \dot{v}^2}{c^3} \gamma^8 \frac{(\gamma \theta)^2}{(1 + \gamma^2 \theta^2)^5}$$



$$\theta_{\max} = \cos^{-1} \left[\frac{1}{3\beta} (\sqrt{1 + 15\beta^2} - 1) \right] \rightarrow \frac{1}{2\gamma}$$

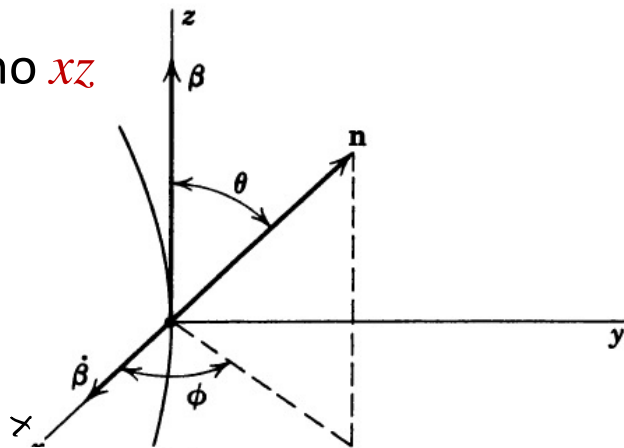
$$\langle \theta^2 \rangle^{1/2} = \frac{1}{\gamma} = \frac{mc^2}{E}$$

Distribuição angular da radiação: velocidade e aceleração perpendiculares

Tomando o movimento circular no plano xz

$$\boldsymbol{\beta} = \beta \hat{\mathbf{z}}$$

$$\dot{\boldsymbol{\beta}} = \dot{\beta} \hat{\mathbf{x}}$$



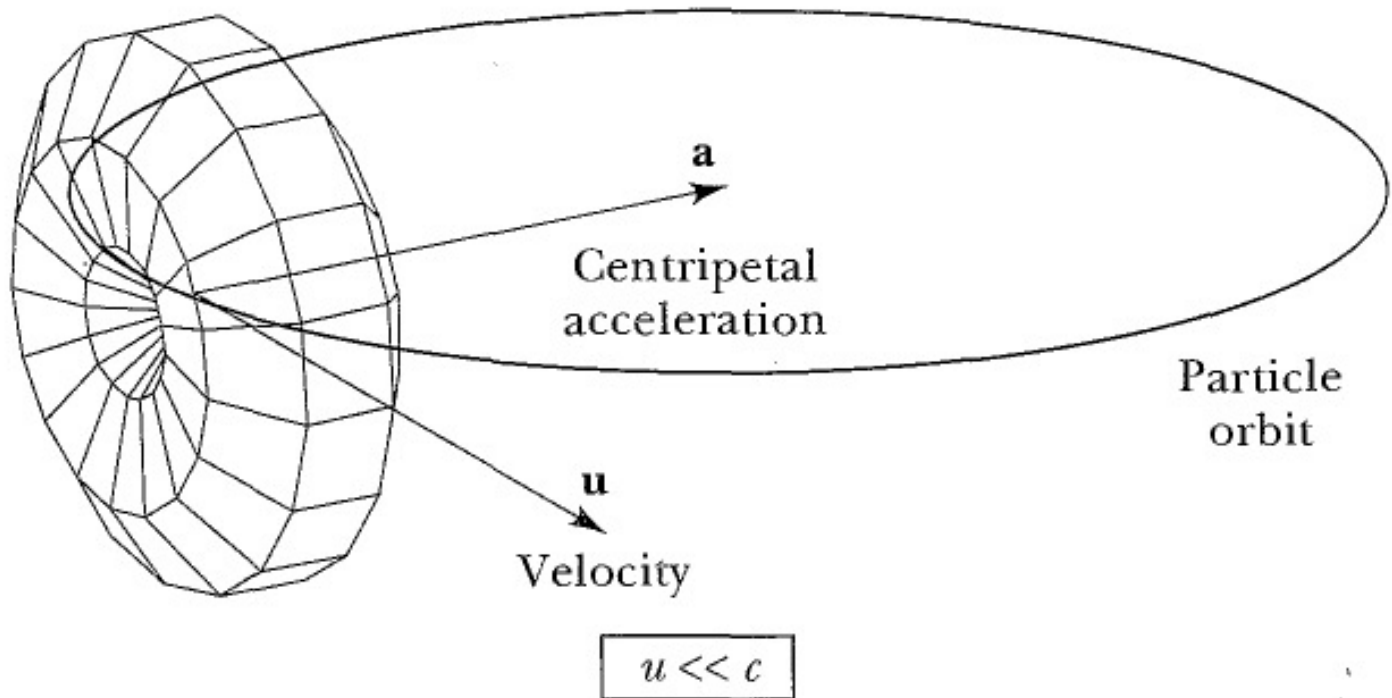
$$\frac{dP(t')}{d\Omega} = \frac{e^2}{4\pi c^3} \frac{|\dot{\mathbf{v}}|^2}{(1 - \beta \cos \theta)^3} \left[1 - \frac{\sin^2 \theta \cos^2 \phi}{\gamma^2 (1 - \beta \cos \theta)^2} \right]$$

$$\gamma \gg 1 \quad \frac{dP(t')}{d\Omega} \simeq \frac{2}{\pi} \frac{e^2}{c^3} \gamma^6 \frac{|\dot{\mathbf{v}}|^2}{(1 + \gamma^2 \theta^2)^3} \left[1 - \frac{4\gamma^2 \theta^2 \cos^2 \phi}{(1 + \gamma^2 \theta^2)^2} \right]$$

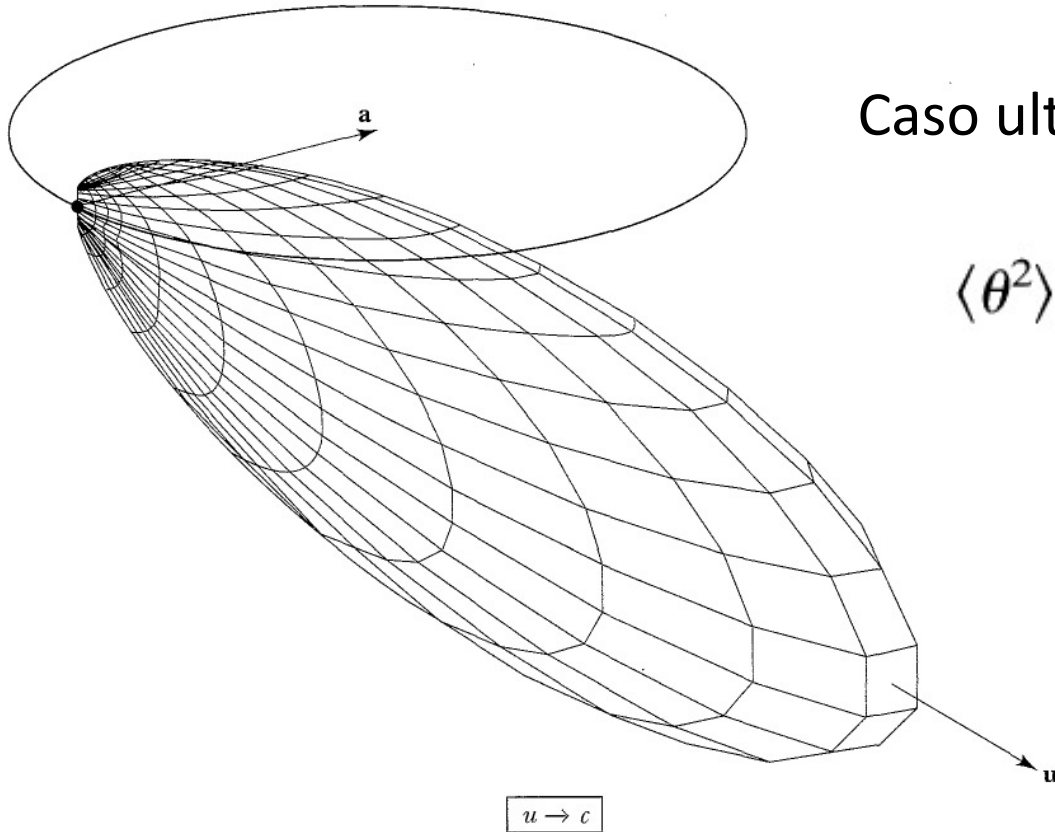
$$\langle \theta^2 \rangle^{1/2} = \frac{1}{\gamma} = \frac{mc^2}{E}$$

Distribuição angular da radiação: velocidade e aceleração perpendiculares

Caso não-relativístico:



Distribuição angular da radiação: velocidade e aceleração perpendiculares

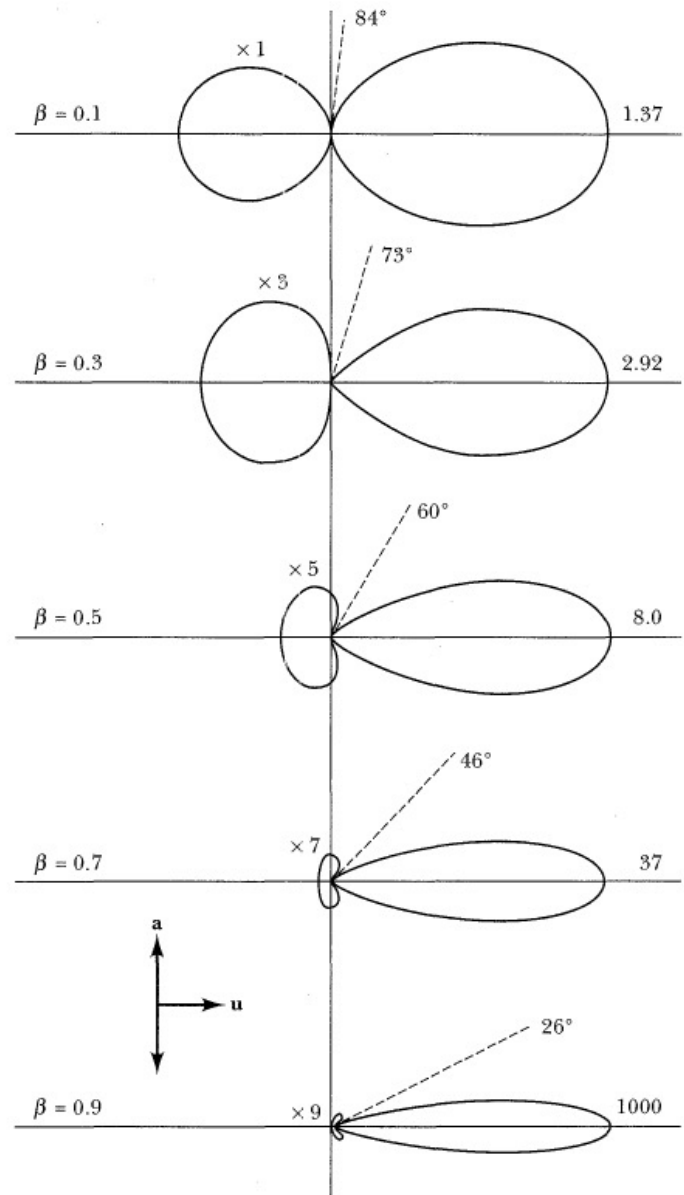


Caso ultra-relativístico:

$$\langle \theta^2 \rangle^{1/2} = \frac{1}{\gamma} = \frac{mc^2}{E}$$

Dependência com a velocidade:

$$\langle \theta^2 \rangle^{1/2} = \frac{1}{\gamma} = \frac{mc^2}{E}$$



Distribuição espectral da radiação

Energia total irradiada por ângulo sólido $d\Omega$ e no intervalo $[\omega, \omega+d\omega]$

$$\frac{d^2 I}{d\omega d\Omega} = \frac{e^2 \omega^2}{4\pi^2 c} \left| \int_{-\infty}^{\infty} \hat{\mathbf{n}} \times (\hat{\mathbf{n}} \times \boldsymbol{\beta}) e^{i\omega[t - \hat{\mathbf{n}} \cdot \mathbf{r}(t)/c]} dt \right|^2$$

